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^{10}Be age datings of Late Würmian erratics from Vorarlberg (Austria) and southern Germany

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Abstract

This study presents a ^{10}Be dataset from Vorarlberg (Austria) and the adjacent Alpine foreland of southern Germany. Out of a total of forty-three samples collected in the field, twenty-five samples of quartz-bearing erratic boulders of Late Würmian age were selected for analyses, measurements and calculations after careful pre-analysis assessment of the sample quality and the potential stratigraphic value. The samples were taken from five localities in Vorarlberg and one locality in southern Germany. The output of fifteen samples passes the threshold in a statistical accuracy assessment. These samples are used in the final determination of the minimum exposure ages. Subsequently, the datings were assessed versus the existing regional morpho- and chronostratigraphic framework, which was erected based on geomorphological and sedimentological data in combination with ^{14}C age datings. The ^{10}Be ages at three localities fit in with the stratigraphic framework, while three are considered slightly young or too young when compared with the existing chronologies. The results are considered promising and justify further ^{10}Be analysis to improve the deglaciation chronology of the area and to find potential calibration sites for cosmogenic surface exposure dating.

Key words: ^{10}Be age datings, erratic boulders, Vorarlberg, Allgäu, mountains and foreland, Late Würmian

1 Introduction

Dating of landforms and erratics based on cosmogenic ^{10}Be data has become an established technique in geomor-

phology and Quaternary stratigraphy (BRAUMANN et al. 2021, 2022; COCKBURN & SUMMERFIELD 2004; IVY-OCHS et al. 2007; KAMLEITNER et al. 2022; PALLÀS et al. 2010). A sampling programme for ^{10}Be age

dating was carried out by us in Vorarlberg and southern Germany in the relatively early days of the application of the technique, i. e. in 2006. Results of the analyses and the calculations

became available in 2011. The findings remained »on the shelf« for reasons beyond the scope of this paper. They are here presented, not in the least in commemoration of our colleague Leo de Graaff, who deceased in December 2021.

In our study area chronological constraints for the Late Würmian are conventionally based on ¹⁴C age datings. Their number, however, is very limited and almost all samples were collected in the foreland (DE JONG et al. 2011; KELLER & KRAYSS 2005b; SPÖTL et al. 2018; KAMLEITNER et al. 2022). Hence, there was and is a need for additional age datings, not only to refine the local and regional stratigraphy, but also to have a means for linking the inferred climate changes with Northern Hemisphere climate proxies, e. g. Greenland Ice Sheet ice-core data (DE JONG et al. 2011).

Geomorphological and glacio-geological research has been carried out in Vorarlberg by the University of Amsterdam since the mid-1950s. The research focus was extended from the primarily glacial-erosive landscape of the mountains of Vorarlberg to the glacial-depositional landscape of the Alpine foreland of southern Germany in the mid-1970s. The investigations resulted in a series of publications on glacial geomorphology and Quaternary stratigraphy. The reader is referred to <https://rfase.org/publications-products/> for a comprehensive overview. ESRI Storymap <https://bit.ly/vorarlberg-storymap> can be consulted for results of the geomorphological mapping.

The deglaciation history of the period following the last major Würmian glacier advance (PREUSSER 2004; HEIRI et al. 2014) is presented in detail by SIMONS (1985) for the eastern Rhine Valley flank and the Walgau region of Vorarlberg. Similarly, DE JONG et al. (1995; <https://rfase.org/466-2/>) review the deglaciation history of the Vordere Bregenzerwald and the adjacent areas of the Rotach and Weissach valleys in Germany. The morphostratigraphic

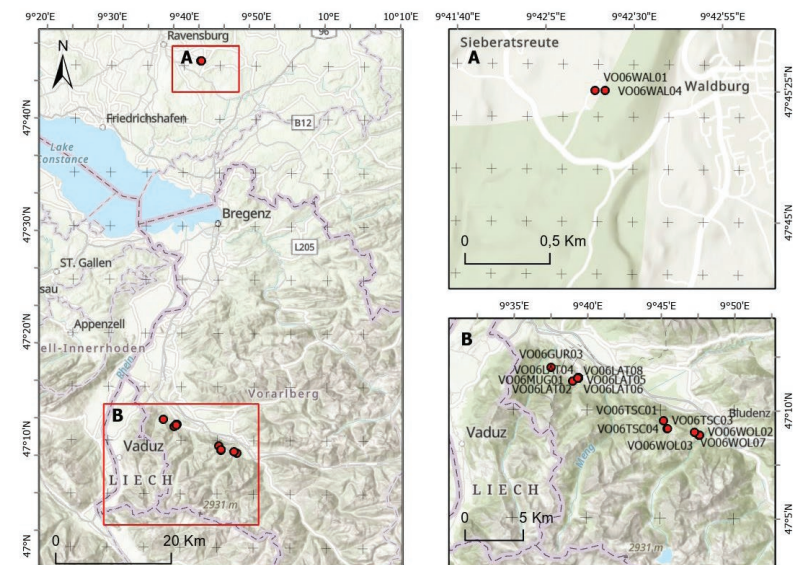


Fig. 1: Sites of samples used in the final determination of the minimum exposure ages. Backdrop maps: »World Topographic Map« and »World Hillshade« from Esri, 2022.

framework of these areas is tied to that erected for the area of the Argen lobe of the Rhine Glacier in the German foreland (DE JONG 1983). SEIJMONSBERGEN (1992) unravels the stratigraphy of the landforms and deposits of the northwestern Rätikon and southern Walgau. A more recent summary of the geomorphological evolution of Vorarlberg is given by DE GRAAFF et al. (2007) and DE JONG et al. (2011). We use the stratigraphic scheme of these publications.

KELLER & KRAYSS (2005a, 2005b) present an overview of their geomorphological and stratigraphic research of the entire area of Rhine-Linth Glacier during several decades. Although there are many differences on a local level, their stratigraphic scheme is similar on a regional scale to that of our paper. HEIRI et al. (2014; see also PREUSSER 2004) review the palaeoclimatic records of the Austrian and Swiss Alps for the period 60-8 ka, and thus present the larger stratigraphic framework for our study. Recently, new age datings have become available in an overview of the entire Rhine and Reuss glacier areas by KAMLEITNER et al. (2022).

2 Location and sample selection

Samples of erratic blocks were selected on glacial and ice-marginal landforms in Vorarlberg and adjacent Allgäu based on local knowledge of the morphogenetic domains, deposits and stratigraphic position, documented in publications and maps (DE GRAAFF et al. 2007; DE JONG 1983; SEIJMONSBERGEN 1992; SIMONS 1985; <https://bit.ly/vorarlberg-storymap>; Fig. 1). Sampling-site selection in the foreland was aided by cartographic information (»erratic boulder«) from Topographic Sheet 1:25,000 8224 Waldburg (released in 1973). Samples were taken from quartz-bearing boulders (metamorphic rocks; $\approx \geq 0.5\text{-}1\text{ m}^3$; Fig. 2). In most cases, sites were selected where a minimum of four boulders could be sampled which were judged to have been exposed at the surface since deposition by the glacier and lying in situ ever since. To avoid underexposure caused by block rotation or till cover, only boulders were sampled which were well-anchored in late-glacial deposits or lying directly on the stable bedrock surface. Boulders sticking out more than 1 m above the ground were favoured. Core samples with a length of 2 cm were cut by electric drilling in the

boulder surface (*Table A in Appendix*). After careful pre-analysis assessment of the sample quality and the potential stratigraphic value twenty-five samples were selected for ^{10}Be analysis and calculation of exposure ages, out of a total of forty-three samples collected in the field. They are from locations in the lower Ill Valley of Vorarlberg (Gurtis = 4 samples; Tschengla = 4 samples; Latz = 7 samples; St. Wolfgang = 6 samples; Muggabill = 1 sample) and from the Waldburg area in the Alpine foreland (Waldburg = 3 samples). See *Table A in Appendix* for the exact coordinates of the twenty-five selected samples.

3 Laboratory treatment, measurements and calculations

Laboratory work was done at the University of Barcelona. Samples were crushed and sieved to extract the 250–1000 μm granulometric quartz fraction. Minerals other than quartz were dissolved by HCl and H_2SiF_6 and remaining quartz grains were cleaned using sequential HF dissolutions to remove any potential atmospheric ^{10}Be (BROWN et al. 1991; KOHL & NISHIZUMI 1992; CERLING & CRAIG 1994). Clean quartz cores were then completely dissolved in HF and spiked with 300 μg of ^9Be carrier (BOURLÈS 1988; BROWN et al. 1992). Beryllium was separated by successive solvent extractions and alkaline precipitations (*Table A in Appendix*; BOURLÈS 1988; BROWN et al. 1992). Measurements of ^{10}Be concentrations were taken at the ASTER AMS facility in Aix-en-Provence, France (*Table A in Appendix*). The measured $^{10}\text{Be}/^9\text{Be}$ ratios were corrected for lab procedural blanks and calibrated with the NIST_27900 beryllium standard (Reference Material 4325, assigned value of $2.79 \pm 0.03 \times 10^{-11}$) and using a ^{10}Be half-life of $1.36 \pm 0.07 \times 10^6$ years (NISHIZUMI et al. 2007). Analytical uncertainties are reported as 1σ and include uncertainties associated with AMS counting statistics, standard uncertainty (certification), and chemi-



Fig. 2: Example of a gneiss erratic at Gurtis.

cal blank measurements. ^{10}Be concentration in quartz samples was calculated by following BALCO (2006), while the surface exposure ages were calculated with the Cronus online calculator v. 3 (BALCO et al. 2008; BALCO 2010; *Table B in Appendix*).

^{10}Be measurements were corrected for topographic shielding, to account for the fact that objects are shielded from cosmic radiation by nearby topographic highs and more distant hills and mountains. Commonly, topographic shielding is manually estimated in the field with measurements of local topography and horizon dip angles. We used digital elevation models (DEMs) to automatically calculate topographic shielding for the samples from Vorarlberg. Following ANDERS (2010), a high-resolution LiDAR (Light Detection And Ranging, 1 m resolution) DEM was used to determine nearby shielding from local topography and a medium-resolution SRTM (Shuttle Radar Topography Mission, resampled to 100 m resolution; FARR et al. 2007) DEM for determining more distant shielding from hills and mountains. The GPS location of each sample location was used as center point from which the maximum horizon dip angles were determined in 360 degrees within 1 km (using the LiDAR DEM) and 50 km (using the SRTM DEM). Shielding corrections were calculated as $\Sigma\alpha_i(1-\text{Sin}^{3.3}\theta)/\Sigma\alpha_i$, where α_i

are the azimuths for the sampling site, and θ are the maximum vertical angles between the sampling and the surface in the α direction.

As the calculations were applied assuming no-erosion conditions and no shielding by snow cover (see KAMLEITNER et al. 2022), the reported exposure ages can be interpreted as minimum ages of the stabilisation of the deposit (erratic). As a general rule (see KAMLEITNER et al. 2022), the oldest exposure age of each deposit was considered to be its minimum age. Where the central value of several samples from the same landform overlap with the one-sigma range of the oldest sample in a group, the landform minimum age was computed as the error-weighted mean of the oldest age cluster. Only sample LAT07 yielded a much older age (with no overlap at 2 sigma ranges with the rest of samples from the same landform). It was interpreted as an outlier owing to inheritance, and was not considered in the minimum age determination. The resulting minimum exposure ages based on the constant production rate model for cosmic-ray exposure (*Table B in Appendix*) are: 16.7 ± 0.9 ka for St. Wolfgang, 16.3 ± 0.6 ka for Tschengla, 18.4 ± 1.8 ka for Gurtis, 16.3 ± 0.8 ka for Latz, 18.8 ± 1.2 ka for Waldburg and 16.2 ± 1.5 ka for Muggabill (*Table 1*).

| ¹⁰ Be (LIFTON et al. 2014) | | | | | | | DE GRAAFF et al. (2007); DE JONG (1983); DE JONG et al. (1995; 2011) | |
|---------------------------------------|--------------------|---------------------------|-------------|--------------|---|---------------------------------------|--|---|
| Sample name | Central value (yr) | External uncertainty (yr) | 1/sigma^2 | mean/sigma^2 | MINIMUM AGE OF LANDFORM (yr) | | | |
| | | | | | Weighted mean | External uncertainty | | |
| MUG01 | 16172 | 1498 | 4,4563E-07 | 7,21E-03 | 16172 | 1498 | Muggabill | |
| TSC02 | 13726 | 1355 | 5,44655E-07 | 7,48E-03 | | | 16275 | 649 |
| TSC03 | 15976 | 1160 | 7,4316E-07 | 1,19E-02 | ≈ RC III; age considered slightly young | | | |
| TSC01 | 16217 | 1110 | 8,1162E-07 | 1,32E-02 | | ≈ end RC II; age considered too young | | |
| TSC04 | 16604 | 1104 | 8,2047E-07 | 1,36E-02 | | | | |
| WOL06 | 13831 | 1568 | 4,06732E-07 | 5,63E-03 | | 16740 | 900 | St. Wolfgang |
| WOL01 | 14555 | 1856 | 2,90298E-07 | 4,23E-03 | ≈ younger than RC IV; age considered reliable | | | |
| WOL04 | 14564 | 2030 | 2,42665E-07 | 3,53E-03 | | | | |
| WOL03 | 16374 | 1245 | 6,4515E-07 | 1,06E-02 | Latz | | | ≈ younger than RC IV; age considered reliable |
| WOL02 | 16675 | 2285 | 1,9153E-07 | 3,19E-03 | | | | |
| WOL07 | 17366 | 1586 | 3,9755E-07 | 6,90E-03 | | | | |
| LAT07 | 23867 | 2337 | 1,83098E-07 | 4,37E-03 | 16344 | | | 779 |
| LAT03 | 13749 | 1735 | 3,32201E-07 | 4,57E-03 | | | | |
| LAT08 | 15408 | 2009 | 2,4777E-07 | 3,82E-03 | | | | |
| LAT04 | 15567 | 1970 | 2,5767E-07 | 4,01E-03 | | | | |
| LAT06 | 16615 | 1738 | 3,3106E-07 | 5,50E-03 | | | | |
| LAT05 | 16712 | 1359 | 5,4145E-07 | 9,05E-03 | Gurtis | ≈ RC II; age considered reliable | | |
| LAT02 | 16874 | 1929 | 2,6874E-07 | 4,53E-03 | | | | |
| GUR04 | 15079 | 1442 | 4,80916E-07 | 7,25E-03 | | | | |
| GUR01 | 15367 | 1401 | 5,09476E-07 | 7,83E-03 | 18396 | 1782 | Waldsburg | |
| GUR02 | 16327 | 1245 | 6,45151E-07 | 1,05E-02 | | | | |
| GUR03 | 18396 | 1782 | 3,1491E-07 | 5,79E-03 | | | | |
| WAL02 | 16273 | 1762 | 3,22098E-07 | 5,24E-03 | 18825 | 1159 | Waldsburg | |
| WAL01 | 18613 | 1729 | 3,3451E-07 | 6,23E-03 | | | | |
| WAL04 | 18998 | 1562 | 4,0986E-07 | 7,79E-03 | | | | |

*Age dating of 20.6 ± 1.7 ka for a boulder at Waldsburg by KAMLEITNER et al. (2022) suggests that 18.8 ± 1.2 ka may be slightly young.
In bold and italics: samples contributing to minimum age calculation

Table 1: Minimum ages – error-weighted means and external uncertainties.

The stratigraphic interpretation is displayed in the right column.

4 Stratigraphic interpretation

The minimum ages of the erratics (*Table 1*) have been assessed versus the position of the corresponding landforms and deposits in our regional morpho- and chronostratigraphic scheme (DE GRAAFF et al. 2007; DE JONG et al. 2011), which evolved from the work of DE JONG (1983) and DE JONG et al. (1995). The Waldsburg ridge of the Alpine foreland consists of ice-marginal and predominantly glaciofluvial landforms and deposits formed during the early phases of glacier recession after the last major Würmian advance of the Rhine Glacier. The landforms and deposits are interpreted to be transitional between Recessional Complex (RC) I and RC II. The age of 18.8 ± 1.2 ka is considered to be slightly young. KAMLEITNER et al. (2022) present ¹⁰Be surface exposure ages for boulders from the same locality at Waldsburg where we collected our samples: 17.9 ± 0.6 ka

(sample RH08) and 20.6 ± 1.7 (sample RH09). Following the rules for handling groups of samples (refer to section Laboratory treatment, measurements and calculations) by combining samples WAL04 and RH09, the minimum exposure age of the landform is 19.7 ± 1.2 ka, which is in line with our chronostratigraphy.

The landforms of Gurtis (SEIJMONSBERGEN et al. 2014: Chapter 5) and Tschengla are among the oldest of a well-developed series of ice-marginal ridges and terraces on the southern flank of the Ill Valley. The age of 16.3 ± 0.6 ka for Tschengla appears too young compared to that of 18.4 ± 1.8 ka for Gurtis. Accepting the latter age dating, these landforms are slightly younger than the above-mentioned landforms and deposits at Waldsburg. They are, then, near-synchronous with RC II.

The alluvial-fan terraces at Latz (SEIJMONSBERGEN et al. 2014: Chapter 5) and the ice-marginal features at St. Wolf-

gang are younger (16.3 ± 0.8 ka and 16.7 ± 0.9 ka, resp.) than the landforms and deposits at Gurtis and Tschengla. In the regional context, they are tentatively interpreted to be younger than the RC IV deposits in the Vorderer Bregenzerwald (DE JONG et al., 1995; <https://rfase.org/466-2/>). The minimum exposure ages fit in with this interpretation.

The ice-marginal terrace on top of which the Muggabill boulder rests (KRIEG & VERHOFSTAD 1986: pages 78-79), is stratigraphically located between the Gurtis and Latz landforms. The minimum exposure age of 16.2 ± 1.5 ka must be considered slightly young.

5 Concluding remarks

Application of the technique of dating landforms and erratics based on cosmogenic ¹⁰Be data has proven useful in constraining chronologically

the existing stratigraphic scheme. Further application in the study area is considered promising and, hence, recommended. KAMLEITNER et al. (2022), though, report that few boulders only out of a large number available in their study qualified for surface exposure dating, thus tempering expectations. Three out of six datings are in line with our stratigraphic scheme. Two ages are slightly young and one too young, the reasons for which remain speculative. Notably useful is the relationship now established over relatively long distance between the foreland cluster of samples of Waldburg and the inner-Alpine samples of the Ill Valley. Recent work by BRAUMANN et al. (2021, 2022) in high regions of the Silvretta Massif of Vorarlberg, in the catchment of the Rhine Glacier complex, documents the presence of Late Glacial and Early Holocene moraines based on geomorphological studies and ^{10}Be surface exposure datings. These findings support the results of our research in terms of applicability of the technique and of the regional stratigraphy.

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7 Authors' contributions

L. W. S. de Graaff, who deceased on December 30th, 2021, initiated the research, prepared and carried out the field work and wrote a first draft of the paper. R. Pallàs Serra and Á. Rodés participated in the field work and are responsible for the laboratory analyses and corresponding calculations. N. S.

Anders carried out the calculations for topographic shielding. A. C. Seijmonsbergen provided general support. M. G. G. De Jong participated in the field work, collated all data and coordinated and wrote the final version of the paper.

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Appendix Table A

Field and laboratory data used to calculate the ^{10}Be exposure ages of erratics from the Ill Valley (Vorarlberg) and in the Alpine foreland (Germany). Names of ^{10}Be samples (standardisation): NIST_Certified. Sample density (2.65 g cm^{-3}) and assumed erosion rate (0 cm yr^{-1}) are similar for all samples.

| Field and laboratory data for all samples analyzed | | | | | | | | |
|--|-------------|----------------------------|-----------------------------|----------------------------|-----------------------|-----------------------------------|---------------------------------------|-----------------------------|
| Sample name | Sample type | Latitude ^a (DD) | Longitude ^a (DD) | Elevation (m) ^c | Sample thickness (cm) | Shielding correction ^b | ^{10}Be content ^d | |
| | | | | | | | atoms g^{-1} | \pm atoms g^{-1} |
| WOL01 | | 47.14722 | 9.79306 | 762 | 2.0 | 0.99060 | 112380 | 12644 |
| WOL02 | | 47.14722 | 9.79361 | 753 (743.1) | 2.0 | 0.99250 | 128574 | 15825 |
| WOL03 | | 47.14722 | 9.79389 | 683 (732.1) | 2.0 | 0.99060 | 118420 | 5623 |
| WOL04 | | 47.14861 | 9.79000 | 782 | 1.5 | 0.99250 | 115120 | 14470 |
| WOL06 | | 47.14944 | 9.78806 | 773 | 2.0 | 0.99250 | 107818 | 10391 |
| WOL07 | | 47.14944 | 9.78806 | 773 (773.6) | 2.0 | 0.99250 | 136434 | 9439 |
| TSC01 | erratics | 47.15806 | 9.75361 | 1165 (1155.2) | 1.0 | 0.99700 | 180240 | 6155 |
| TSC02 | Ill Valley | 47.14972 | 9.76194 | 1210 | 2.0 | 0.99690 | 156057 | 12293 |
| TSC03 | | 47.15194 | 9.75833 | 1228 (1227.2) | 2.5 | 0.99716 | 184885 | 7732 |
| TSC04 | | 47.15167 | 9.75778 | 1215 (1222.6) | 2.3 | 0.99700 | 190577 | 5716 |
| GUR01 | | 47.19528 | 9.63556 | 998 | 2.0 | 0.98991 | 145916 | 10070 |
| GUR02 | | 47.19528 | 9.63444 | 1005 | 1.0 | 0.98161 | 156175 | 7470 |
| GUR03 | | 47.19861 | 9.62583 | 977 (973.7) | 2.5 | 0.99194 | 172090 | 13123 |
| GUR04 | | 47.19833 | 9.62583 | 984 | 1.3 | 0.99194 | 142467 | 10653 |
| MUG01 | | 47.18806 | 9.65028 | 879 (873.7) | 2.7 | 0.98322 | 137032 | 9717 |
| LAT02 | | 47.19000 | 9.65639 | 765 (734.8) | 1.5 | 0.98190 | 130715 | 12725 |
| LAT03 | | 47.19028 | 9.65611 | 748 | 2.0 | 0.98691 | 104280 | 11579 |
| LAT04 | | 47.19028 | 9.65611 | 760 (737.2) | 1.0 | 0.98691 | 120943 | 13475 |
| LAT05 | | 47.19028 | 9.65611 | 751 (737.5) | 1.6 | 0.98691 | 128405 | 7126 |
| LAT06 | | 47.19056 | 9.65667 | 752 (730.2) | 1.3 | 0.98691 | 128086 | 10998 |
| LAT07 | | 47.19056 | 9.65667 | 744 | 2.7 | 0.98691 | 183028 | 14174 |
| LAT08 | | 47.19056 | 9.65639 | 740 (732.2) | 2.0 | 0.98190 | 116053 | 13424 |
| WAL01 | erratics | 47.75694 | 9.70528 | 752 (745) | 1.5 | 100.000 | 146764 | 10447 |
| WAL02 | Alpine | 47.75722 | 9.70500 | 741 | 2.7 | 0.99987 | 125362 | 11314 |
| WAL04 | foreland | 47.75694 | 9.70611 | 726 | 2.1 | 0.98589 | 143746 | 8150 |

^a With reference to WGS84 datum

^b Assuming exponent $x = 2.3$ in the cosmic ray flux equation : $\cos x$ (zenith angle)

^c Values between parentheses derived from Vorarlberg DEM-2017 (samples used in minimum age calculation) or adjusted using topographic map (WAL01)

^d Blank values used (ratio $^{10}\text{Be}/^9\text{Be} : 2.41927\text{E}^{-15}$)

Appendix Table B

Exposure ages for constant production rate model. Output from the online exposure age calculator formerly known as the CRONUS-Earth online exposure age calculator v. 3 (BALCO et al. 2008), wrapper: 3.0.2; get_age: 3.0.2; muons: 1A, alpha = 1; validate: validate_v3_input.m - 3.0; consts: 2020-08-26 .

| Sample name | St (STONE 2000) | | | Lm (LAL 1991) | | | LSDn (LIFTON et al. 2014) | | |
|-------------|-----------------|-------------|-------------|---------------|-------------|-------------|---------------------------|-------------|-------------|
| | Age (yr) | Interr (yr) | Exterr (yr) | Age (yr) | Interr (yr) | Exterr (yr) | Age (yr) | Interr (yr) | Exterr (yr) |
| WOL01 | 14480 | 1635 | 1998 | 14183 | 1601 | 1925 | 14555 | 1644 | 1856 |
| WOL02 | 16670 | 2060 | 2448 | 16267 | 2010 | 2354 | 16675 | 2061 | 2285 |
| WOL03 | 16319 | 778 | 1510 | 15931 | 759 | 1420 | 16374 | 781 | 1245 |
| WOL04 | 14500 | 1829 | 2160 | 14201 | 1791 | 2086 | 14564 | 1837 | 2030 |
| WOL06 | 13736 | 1328 | 1718 | 13483 | 1304 | 1652 | 13831 | 1338 | 1568 |
| WOL07 | 17397 | 1209 | 1835 | 16960 | 1178 | 1738 | 17366 | 1207 | 1586 |
| TSC01 | 16466 | 565 | 1423 | 16069 | 551 | 1330 | 16217 | 556 | 1110 |
| TSC02 | 13861 | 1096 | 1552 | 13607 | 1075 | 1485 | 13726 | 1085 | 1355 |
| TSC03 | 16261 | 683 | 1459 | 15870 | 666 | 1368 | 15976 | 671 | 1160 |
| TSC04 | 16913 | 509 | 1435 | 16489 | 497 | 1337 | 16604 | 500 | 1104 |
| GUR01 | 15475 | 1072 | 1630 | 15127 | 1048 | 1548 | 15367 | 1065 | 1401 |
| GUR02 | 16476 | 791 | 1528 | 16081 | 772 | 1436 | 16327 | 784 | 1245 |
| GUR03 | 18618 | 1426 | 2054 | 18111 | 1387 | 1946 | 18396 | 1409 | 1782 |
| GUR04 | 15164 | 1138 | 1656 | 14837 | 1114 | 1577 | 15079 | 1132 | 1442 |
| MUG01 | 16228 | 1155 | 1730 | 15842 | 1128 | 1642 | 16172 | 1151 | 1498 |
| LAT02 | 16882 | 1650 | 2125 | 16468 | 1610 | 2032 | 16874 | 1650 | 1929 |
| LAT03 | 13638 | 1519 | 1865 | 13391 | 1492 | 1800 | 13749 | 1532 | 1735 |
| LAT04 | 15539 | 1738 | 2131 | 15188 | 1699 | 2048 | 15567 | 1741 | 1970 |
| LAT05 | 16709 | 931 | 1620 | 16305 | 909 | 1527 | 16712 | 931 | 1359 |
| LAT06 | 16612 | 1432 | 1946 | 16214 | 1398 | 1856 | 16615 | 1433 | 1738 |
| LAT07 | 24218 | 1887 | 2695 | 23375 | 1821 | 2535 | 23867 | 1859 | 2337 |
| LAT08 | 15364 | 1784 | 2160 | 15024 | 1744 | 2079 | 15408 | 1789 | 2009 |
| WAL01 | 18703 | 1338 | 1998 | 18230 | 1304 | 1894 | 18613 | 1331 | 1729 |
| WAL02 | 16275 | 1475 | 1960 | 15916 | 1442 | 1875 | 16273 | 1475 | 1762 |
| WAL04 | 19086 | 1087 | 1865 | 18593 | 1059 | 1756 | 18998 | 1082 | 1562 |

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